

Infrared emission in metal-halide arc lamps

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HID lamps – IR losses

- Understand and control the infrared (IR) losses in commercial high-intensity discharge (HID) lamps
- Determine the total IR power output up to the fused silica limit ($\sim 3.5\mu\text{m}$)
- Identify the major contributions to losses in the IR
 - Initially Osram Sylvania 250W Metalarc lamp (Na Sc chemistry)
 - Custom Philips HID lamp (rare earth chemistry)

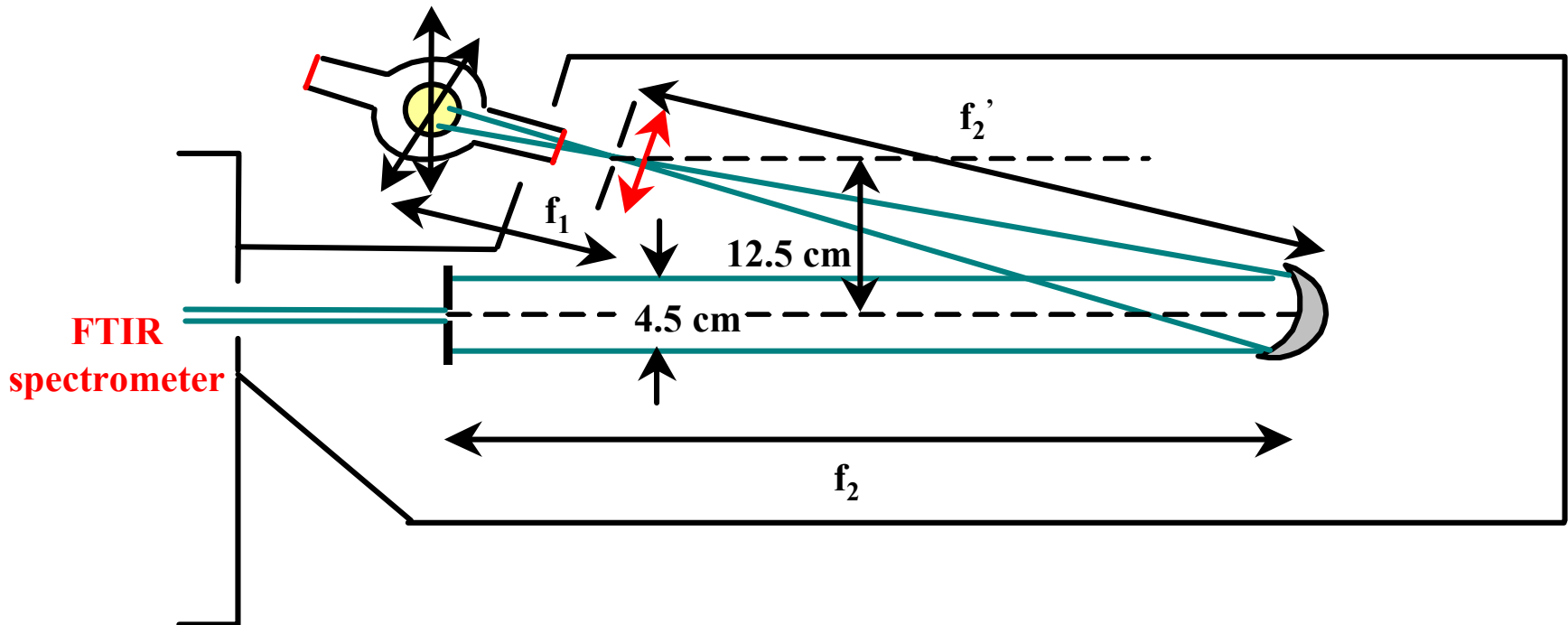


The Lamp

- Osram Sylvania '250W Metalarc' lamp
 - Metal-halide lamp
 - Scandium, Sodium, Caesium, Thorium
- Visible emission spectrum dominated by Sc I
- Lamp chamber is anodized Al with MgF₂ window to allow through the IR
 - chamber pressure of ~15 mTorr



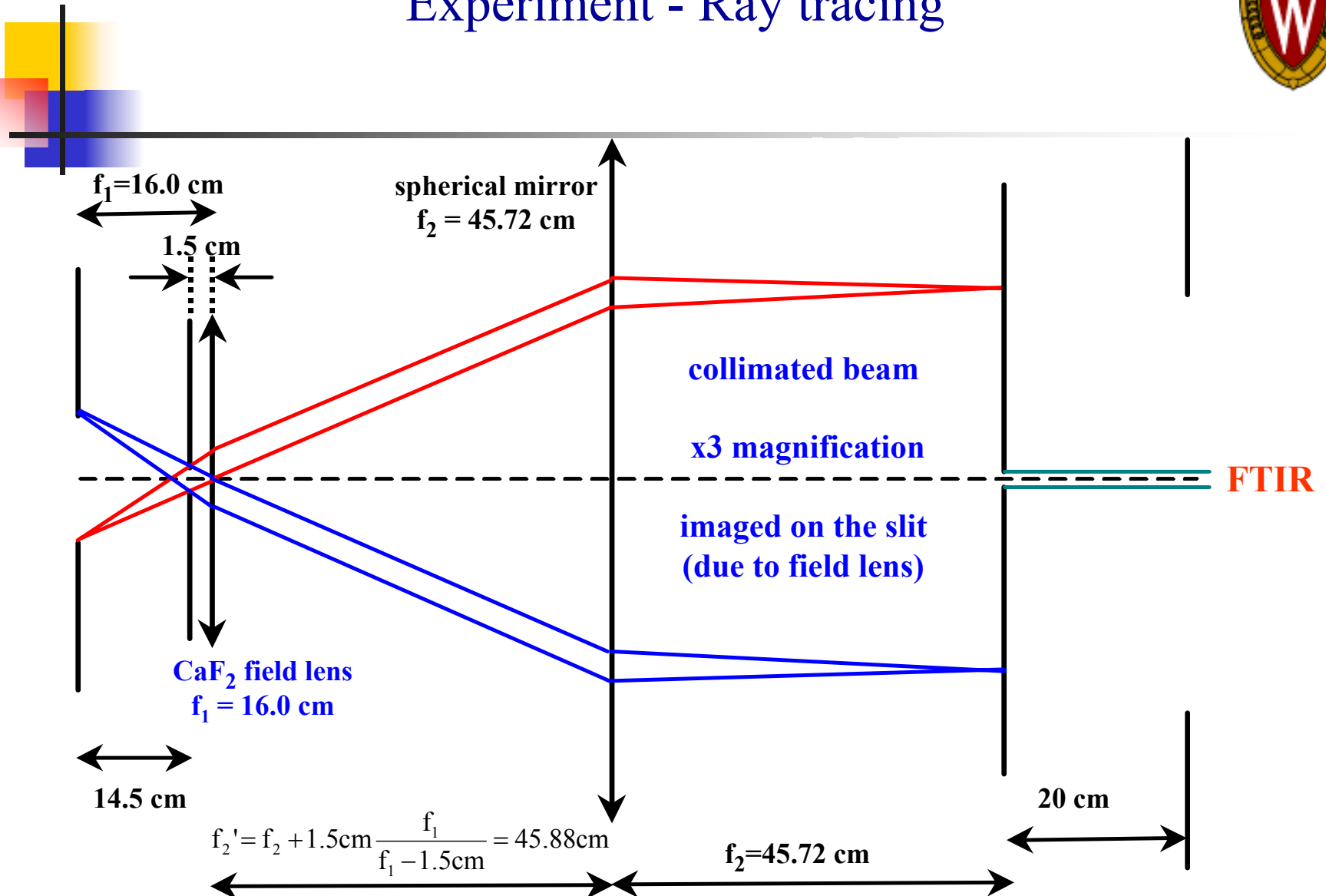
Experiment - optical set-up



external optics encased to ensure dry, CO₂-free air



Experiment - Ray tracing





Experiment - spatial resolution

- Field lens is used to image the arc lamp onto the slit
 - provides excellent spatial resolution
- Lamp is translated on a X-Z translation stage to allow measurements of emission for both radial and longitudinal information
- Impact parameter is unaffected by the arc tube in emission measurements

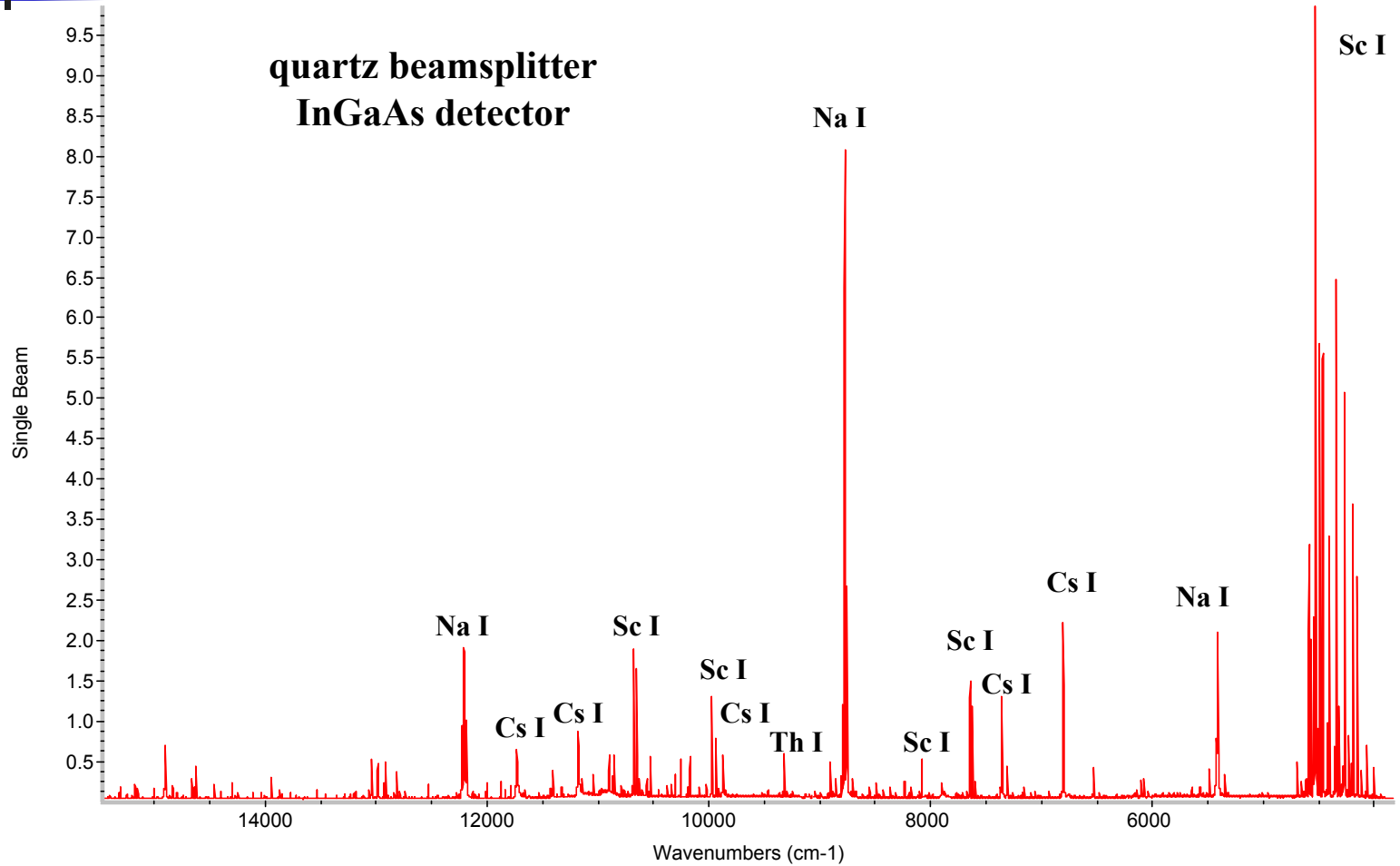
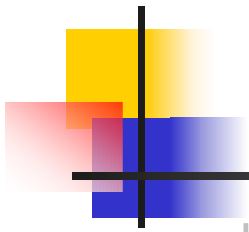


Experimental measurements

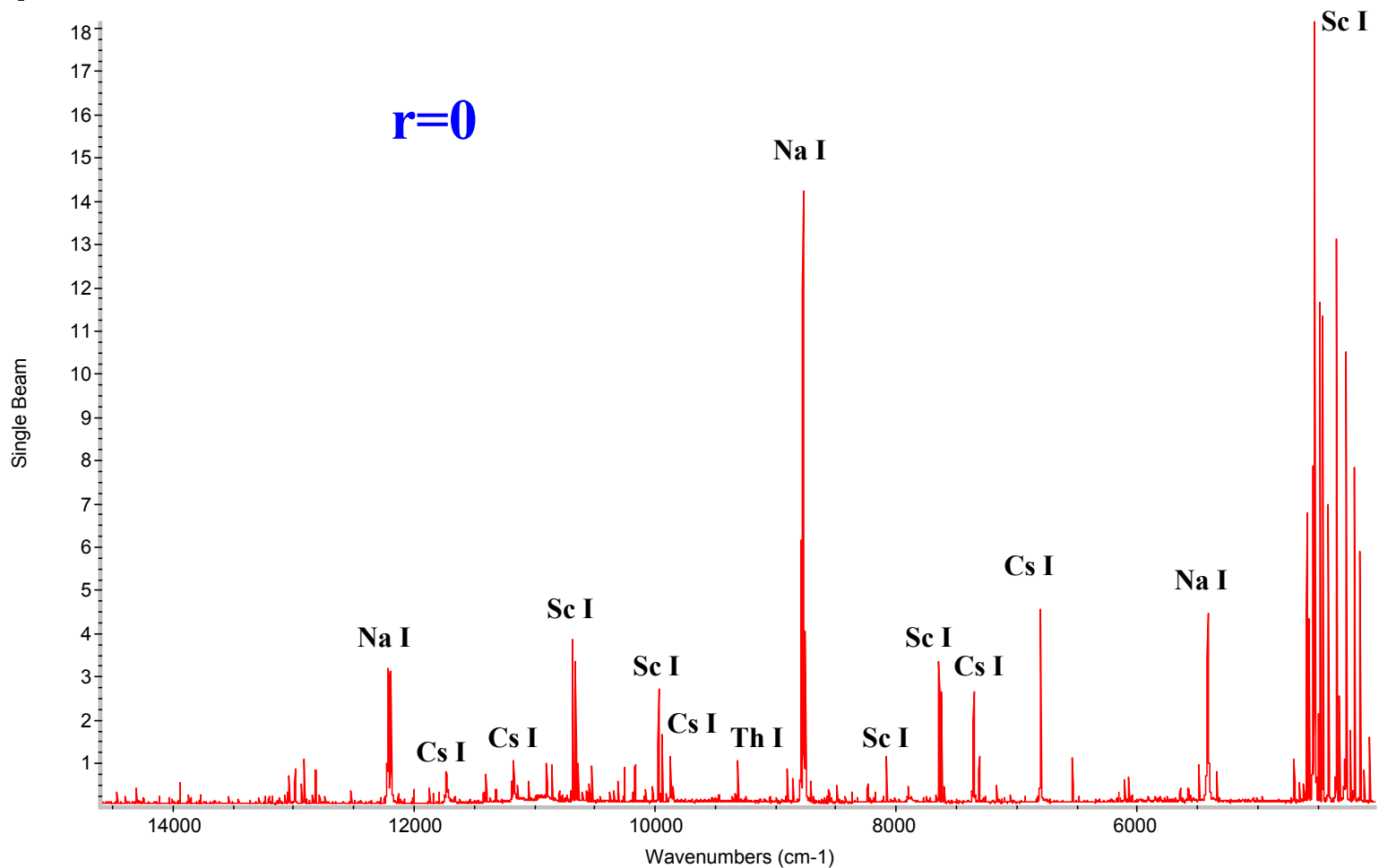
- Fourier transform infrared (FTIR) spectrometer
 - $\sigma = 25,000\text{-}100\text{ cm}^{-1}$
 - MgF_2 window limits the experiment to $1,000\text{ cm}^{-1}$
 - ultimately replace with CaF_2 to measure further into the IR (blackbody radiation)
- Spatially-averaged measurements in the visible, near IR and mid IR
- Spatially-resolved measurements in the near IR
- External optics encased to ensure dry, CO_2 -free air



Spatially-averaged near IR emission (uncorrected spectrum)

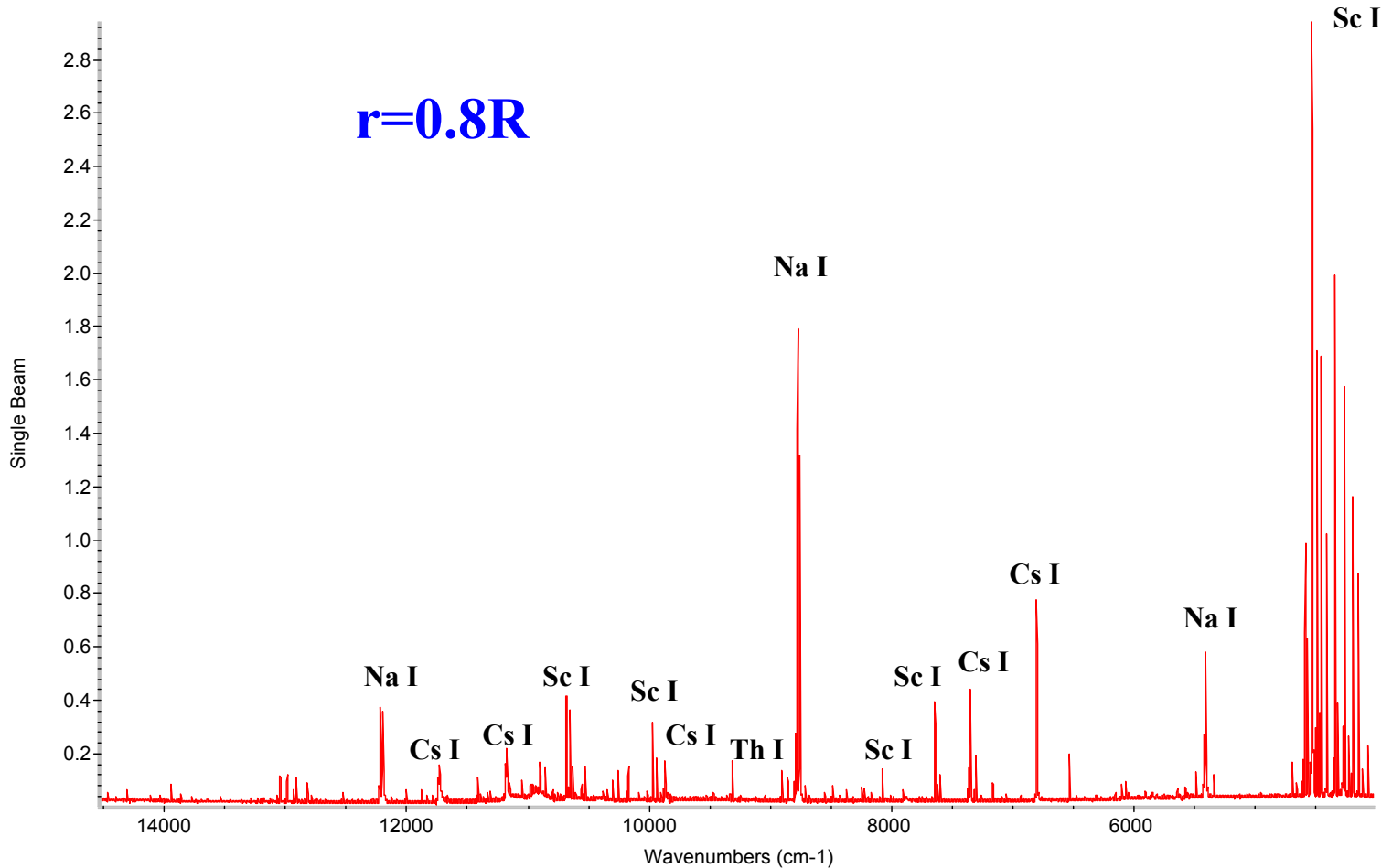


Spatially-resolved near IR measurements (uncorrected spectrum)





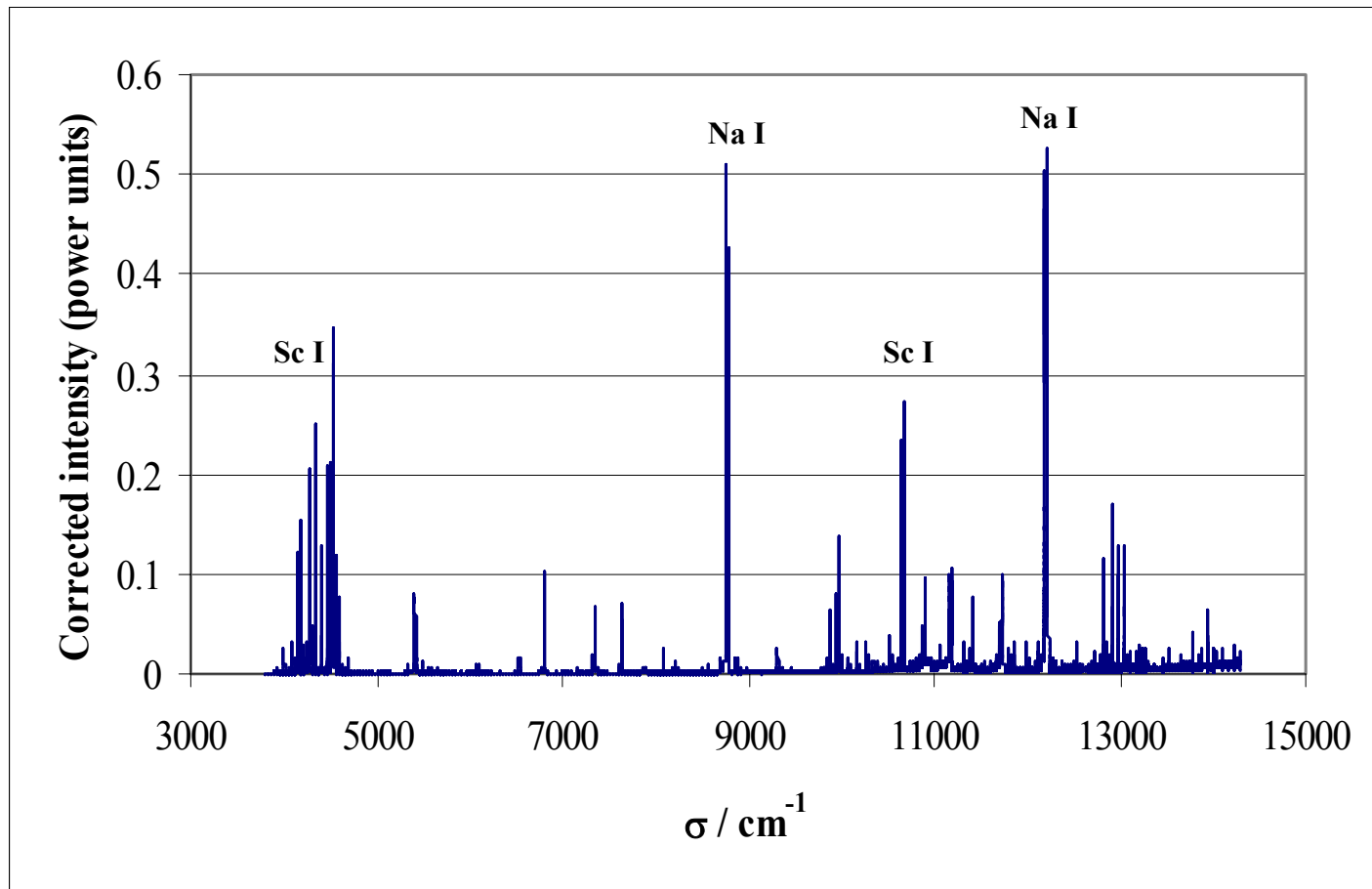
Spatially-resolved near IR measurements (uncorrected spectrum)



Corrected spatially-resolved near IR measurements

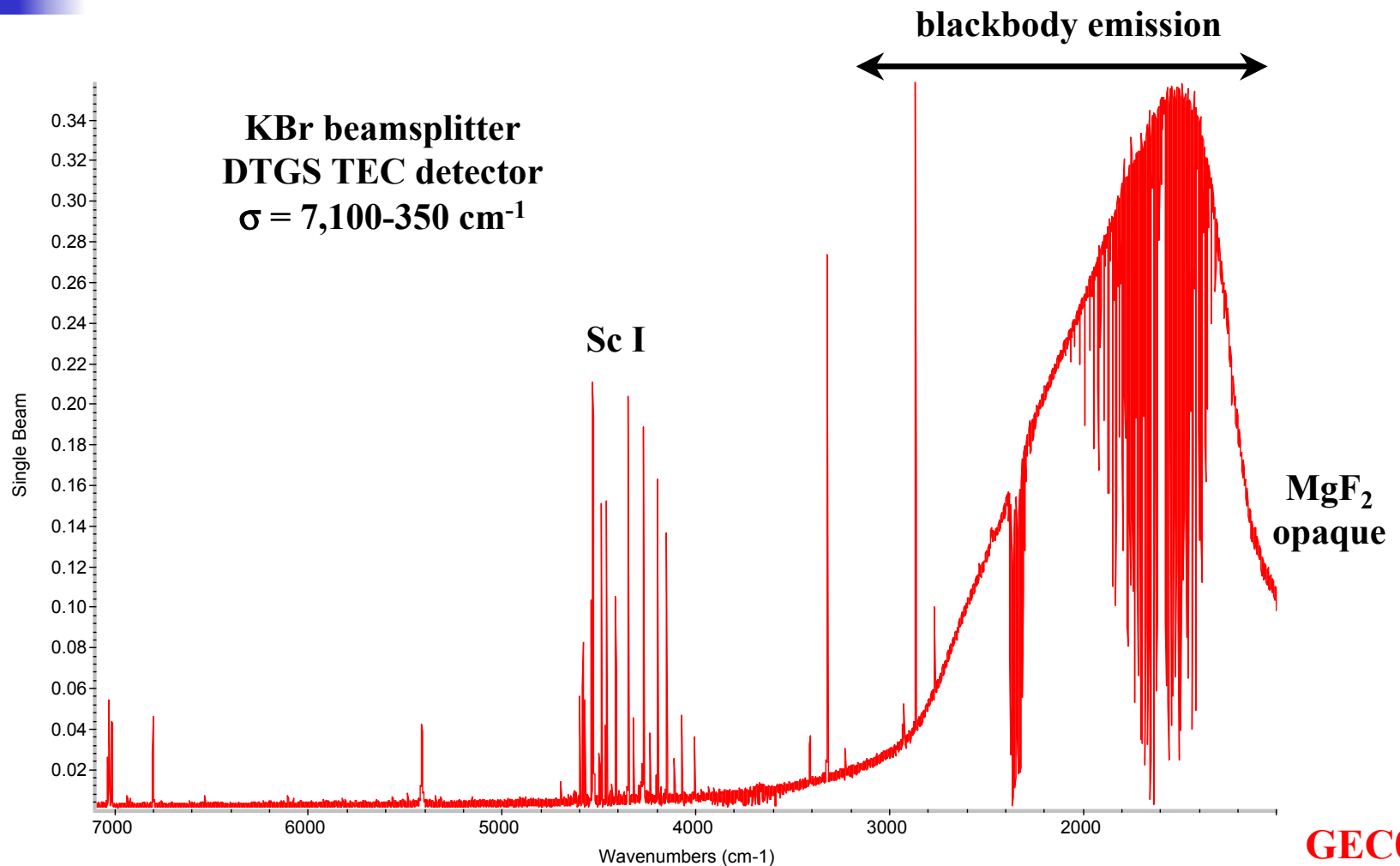


$r=0$



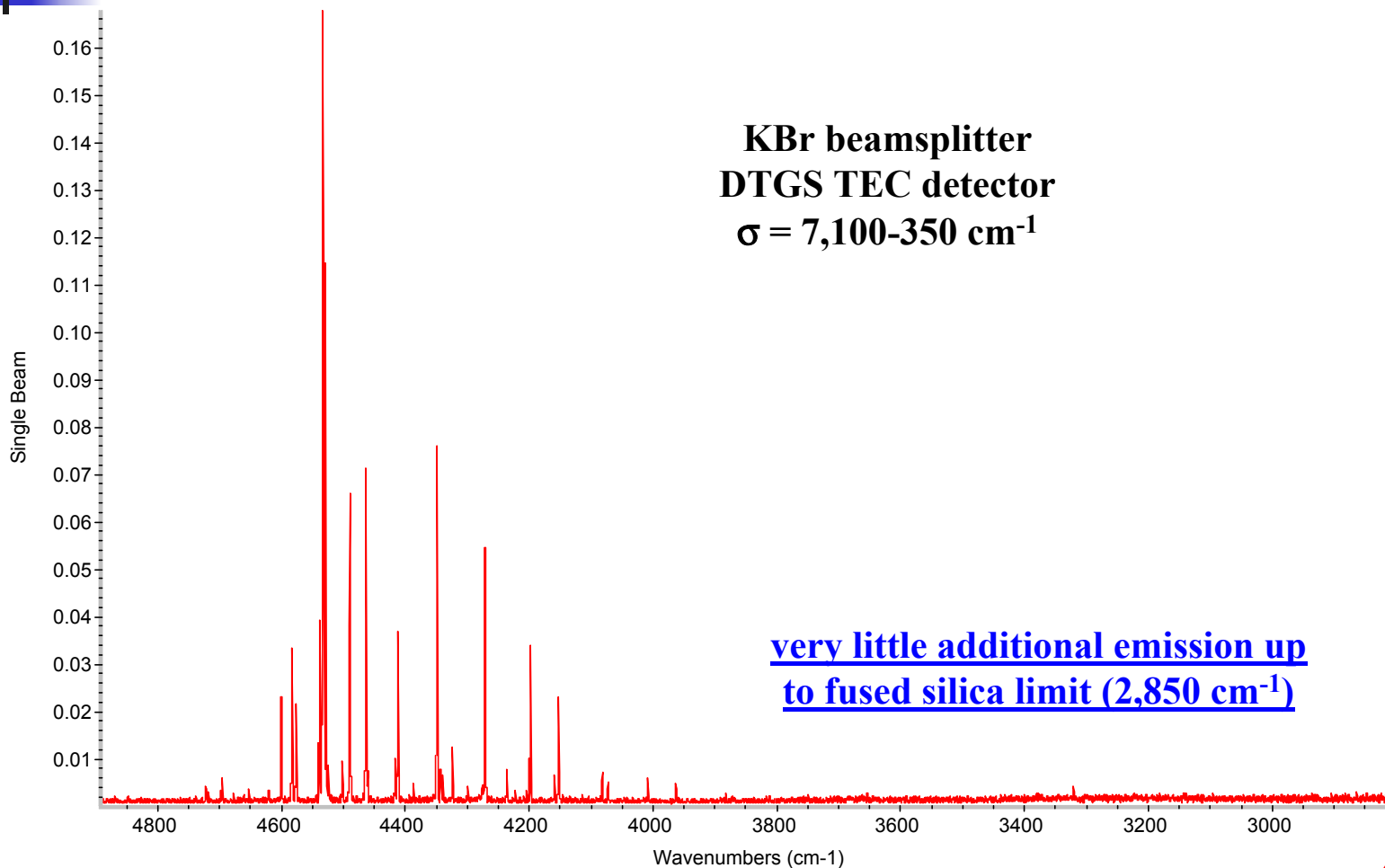


Spatially-averaged mid IR, $\sigma = 7,100-1,000\text{cm}^{-1}$





Beyond 4,000 cm^{-1}



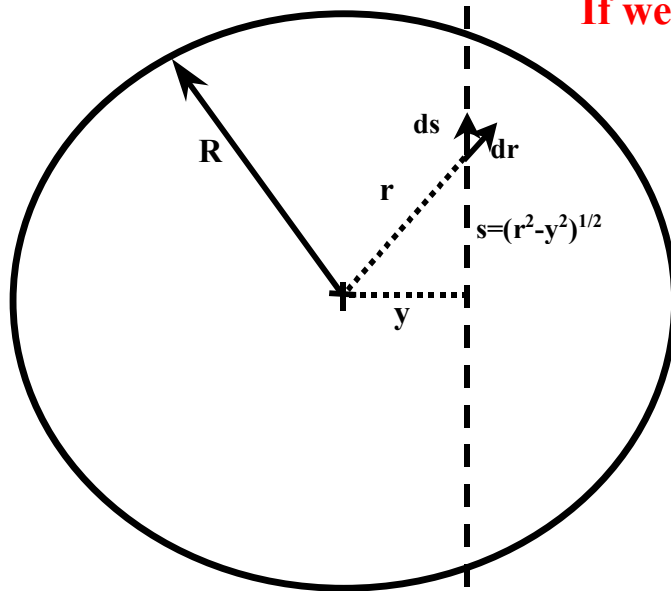


IR power loss

- Near IR losses dominated by Na I emission – we require accurate knowledge of the
 - density of the ground state of Na
 - absorption measurements using the synchrotron radiation source
 - temperature profile, assuming LTE
 - Boltzmann analysis of the absorption data and emission spectra
 - transition probabilities for the important IR transitions



Spatially-resolved emission: Abel Inversion



If we assume that the excited-state densities are described by a Gaussian-like function

which acts to force to zero the values at the wall

$$n(r) = \frac{a \exp\left\{-b \frac{r^2}{1-r^2}\right\}}{(1-r^2)^{3/2}}$$

where r is the relative radius (i.e. normalized to R)

Hence the intensity measured along a line-of-sight, whose point of closest approach is y .

$$I(y) = 2 \int_{s=0}^{\sqrt{1-y^2}} n(r) ds \longrightarrow I(y) = \frac{\sqrt{\pi}}{\sqrt{1-y^2}} \frac{a \exp\left\{-b \frac{y^2}{1-y^2}\right\}}{\sqrt{b}}$$

Spatially-resolved emission: Abel Inversion



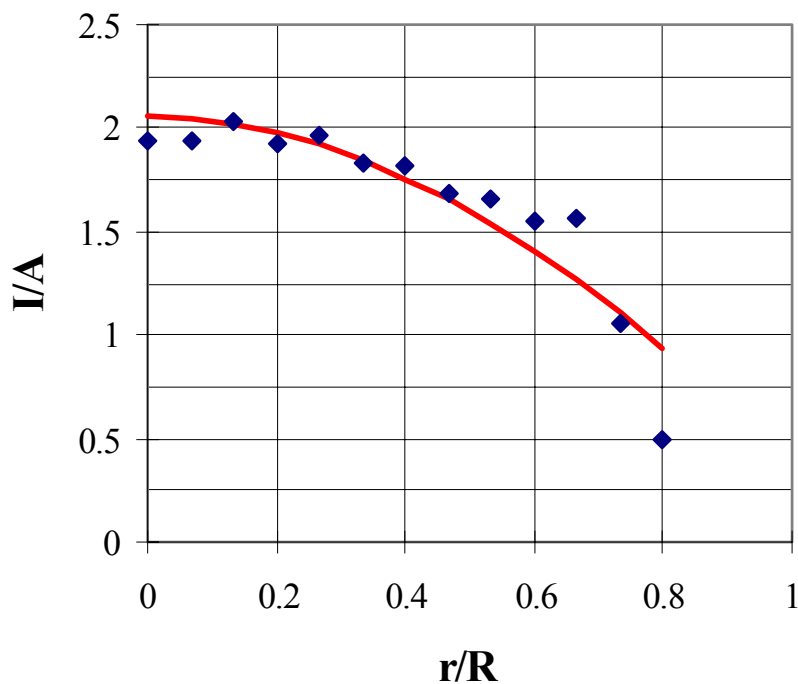
The intensity measured along a line-of-sight, whose point of closest approach is y .

$$I(y) = \frac{\sqrt{\pi}}{\sqrt{1-y^2}} \frac{a \exp\left\{-b \frac{y^2}{1-y^2}\right\}}{\sqrt{b}}$$

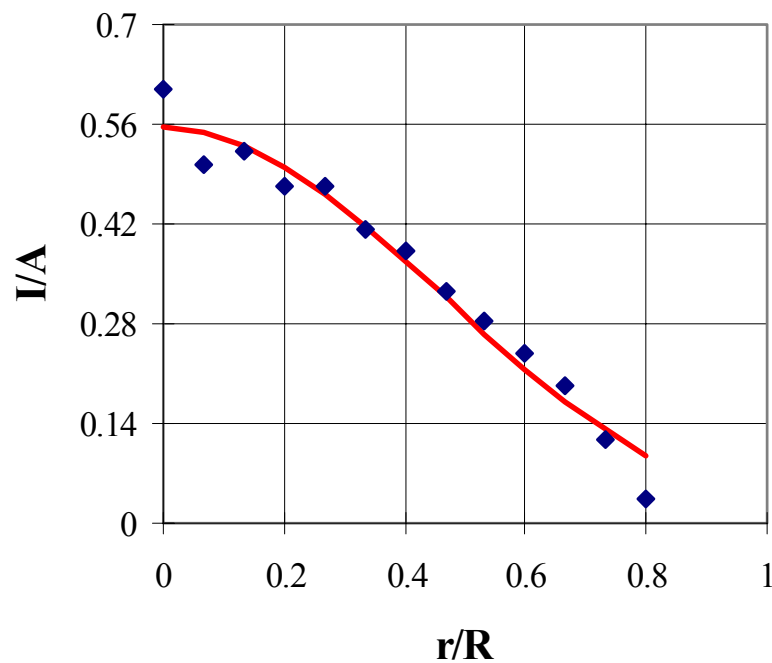
We may then fit this equation to the observed data as a function of impact parameter to obtain numerical values of the parameters a and b .

This gives us $n(r)$

Fitted $I(y)$ assuming a Gaussian profile for the excited-state densities



$$E_k = 16,022 \text{ cm}^{-1}$$
$$a = 1.235, b = 0.5992$$



$$E_j = 21,086 \text{ cm}^{-1}$$
$$a = 0.5042, b = 2.463$$



IR power loss

The total power radiated on a given kJ transition is given by

$$P(\lambda_{kJ}) = \int_V n_k(r, l) A_{kJ} hc \sigma_{kJ} dV$$

which can be expressed in terms of $n_0(r)$ and $T(r)$, where A_{kJ} is known

$$P(\lambda_{kJ}) = A_{kJ} hc \sigma_{kJ} l \int_{r=0}^{0.75\text{cm}} \frac{g_k}{g_0} n_0(r) \exp\left\{-\frac{E_k}{kT(r)}\right\} 2\pi r dr$$

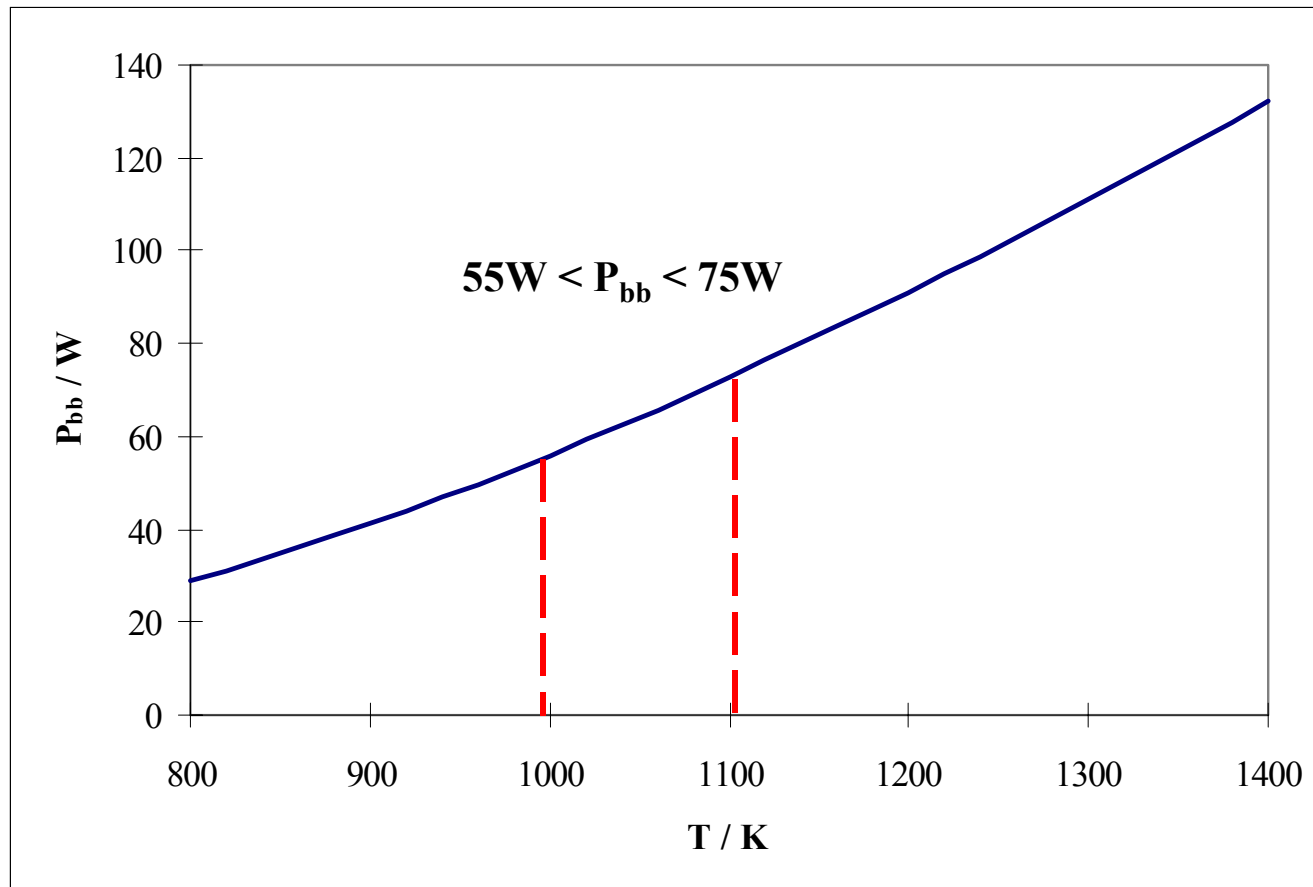


Total IR power loss

- From absorption/emission measurements:
 - total near IR atomic and molecular losses ~ 10-15 W
 - the 3p-3d Na I lines contribute ~ 1 W
 - the 3p-4s Na I lines contribute ~ 3 W
- From thermopile measurements:
 - total near IR atomic and molecular losses ~ 60 W
 - blackbody losses ~ 55-60 W
 - visible emission ~ 85 W
 - UV losses ~ 45 W



Power balance - blackbody radiation



$e = 0$ for $\lambda < 3.5 \mu m$, $e \sim 1$ for $\lambda > 3.5 \mu m$



Conclusions and future work

- Na I dominates atomic emission in the near IR
- Discrepancy between calculated near IR power loss by
 - thermopile measurements (~ 60 W)
 - results calibrated to absorption measurements (~ 10 - 15 W)
- Blackbody radiation contributes a further 60 W to the IR losses

Planned future work

- Resolve the discrepancy
- Extend techniques to custom research lamps
- Compare the IR power / lumen output between Scandium-based lamps and rare-earth lamps



Acknowledgments

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