

EXAM 1 REVIEW

Chapter 1:

- Newton's Laws:
 - 1) In the absence of forces, particles move at constant \vec{v}
 - 2) $\vec{F} = m\vec{a}$
 - 3) $\vec{F}_{12} = -\vec{F}_{21}$
- An inertial frame is any frame of reference in which Law 1 holds.
- Momentum Conservation: For any system of particles $\frac{d}{dt}\vec{P}_{\text{tot}} = \vec{F}_{\text{ext}}$
- General solution method for time dependent forces, $F = F(t)$ – write $m\frac{dv}{dt} = F(t)$, rearrange to separate variables, and integrate:

$$v(t) = v_0 + \frac{1}{m} \int_0^t F(t') dt'$$

To find $x(t)$, integrate one more time.

- Polar coordinates:

$$\vec{a} = (\ddot{r} - r\dot{\phi}^2)\hat{r} + (r\ddot{\phi} + 2\dot{r}\dot{\phi})\hat{\phi}$$

Chapter 2:

- Drag Forces: $\vec{F} = -f(v)\hat{v}$, where $f(v) = bv + cv^2$
- Projectile motion with linear drag:
 - Horizontal motion: $m\dot{v}_x = -bv_x$ – linear, first order, homogenous differential equation – easily solved for $v(t)$ with an exponential function
 - Vertical motion: $m\dot{v}_y = mg - bv_y$ – same equation except inhomogenous – find particular solution and combine with general solution of homogenous equation
- General solution method for velocity dependent forces – write $m\frac{dv}{dt} = F(v)$, rearrange to separate variables, and integrate:

$$t = m \int_{v_0}^v \frac{dv}{F(v)}$$

Solve for $v(t)$ by algebra. Then integrate one more time to find $x(t)$.

- Problems with $F = F(x)$ can be solved for $v(x)$ by writing $m\frac{dv}{dt} = mv\frac{dv}{dx}$. Separate variables and integrate:

$$m \int_{v_0}^v v dv = \int_{x_0}^x F(x) dx$$

This works for velocity dependent forces as well, with $F(v)$ in the v integral.

- Motion in a uniform magnetic field:

$$v_x = v_{x0} \cos \omega t + v_{y0} \sin \omega t \quad v_y = v_{y0} \cos \omega t - v_{x0} \sin \omega t$$

where $\omega = qB/m$. This is circular motion with $R = mv/qB$.

- Motion in crossed electric/magnetic field: Circular motion (as above) superimposed on uniform drift \perp to both \vec{E} and \vec{B} .

Chapter 3:

- Rocket equation: $m \frac{dv}{dt} = -v_{\text{ex}} \frac{dm}{dt} + F_{\text{ext}}$
- Center of Mass: $\vec{R} = \frac{1}{M} \sum_{\alpha} m_{\alpha} \vec{r}_{\alpha}$. This leads to $\vec{P}_{\text{tot}} = M \frac{d}{dt} \vec{R}$ and $\vec{F}_{\text{ext}} = M \frac{d^2}{dt^2} \vec{R}$.
- Angular Momentum: Single particle: $\vec{\Gamma} = \frac{d}{dt} \vec{\ell}$ For a system of particles: $\vec{\Gamma}_{\text{ext}} = \frac{d}{dt} \vec{L}$
- Rotation about an axis: $L = I\omega$ where $I = \sum_{\alpha} m_{\alpha} \rho_{\alpha}^2$
- Parallel axis theorem: $I = I_{\text{cm}} + M\rho^2$
- Two-body collisions – apply energy and momentum conservation
- CM Frame of reference – moves along with CM of the system $\vec{v}_{\alpha} = \vec{v}'_{\alpha} + \vec{v}_{\text{cm}}$
Total momentum in the cm frame is zero.

Chapter 4:

- Work energy theorem: Gain in kinetic energy = work done on m -- $W = \int \vec{F} \cdot d\vec{s}$
- Potential energy: $U(\vec{r}) = - \int_{\vec{r}_s}^{\vec{r}} \vec{F}(\vec{r}) \cdot d\vec{s}$ $\vec{F} = -\vec{\nabla}U$
- Energy ($T + U$) is conserved (and definition of U makes sense) if $\vec{\nabla} \times \vec{F} = 0$
- Problem solving:
 $T + U = E$ (constant determined from initial conditions) gives v as a function of position.
Constraint forces have no effect on E
 $\frac{d}{dt}[T + U] = 0$ may give a useful equation of motion
- Central forces are always conservative: $U(r) = - \int_{r_s}^r F(r) dr$
- Two-body system with no external forces and $\vec{F}_{12} = -\vec{F}_{21} = F(r)\hat{r}$ (where $\vec{r} = \vec{r}_1 - \vec{r}_2$)
 $E = T_1 + T_2 + U(r)$ is conserved
Internal motion given by $\vec{F} = \mu \frac{d^2}{dt^2} \vec{r}$ where $\mu = \frac{m_1 m_2}{m_1 + m_2}$