

DISCUSSION 4/15/04

Review for Exam

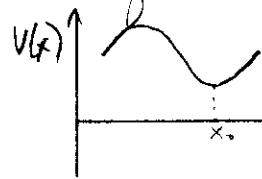
CHAPTER 5  $\Rightarrow$  Oscillations

Beginning with SHM  $F = m\ddot{x} = -kx$

$\Rightarrow x(t) = A \cos(\omega t - \phi)$

$\omega = \sqrt{\frac{k}{m}}$

$\Rightarrow$  Good approx for oscillations about any stable equilibrium pt. if amplitude is small



• Next add a linear damping force

$F = -bv \Rightarrow$

$m\ddot{x} + b\dot{x} + kx = 0$

OR

$Dx(t) = 0$  where  $D = m \frac{d^2}{dt^2} + b \frac{d}{dt} + m$

Solutions:

Strong damping:  $\beta = \frac{b}{2m} > \omega_0$

$x(t) = A_1 e^{-\alpha_1 t} + A_2 e^{-\alpha_2 t}$

$\alpha = \beta \pm [\beta^2 - \omega_0^2]^{\frac{1}{2}}$

Weak damping  $\beta < \omega_0$

$x(t) = A e^{-\beta t} \cos(\omega_1 t - \phi)$

$\omega_1 = [\omega_0^2 - \beta^2]^{\frac{1}{2}}$

Critical

$\beta = \omega_0$

$x(t) = A e^{-\beta t} + B t e^{-\beta t}$

• Forced H.O.

$m\ddot{x} + b\dot{x} + kx = F(t)$

$F(t) = f_0 \cos \omega t$

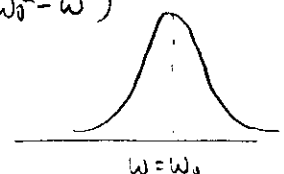
$\Rightarrow x(t) = \underbrace{B \cos(\omega_1 t - \phi)}_{\text{transient}} e^{-\beta t} + A \cos(\omega t - \delta)$

where

$A = \frac{f_0}{[(\omega_0^2 - \omega^2) + (2\beta\omega)^2]^{\frac{1}{2}}}$

$\tan \delta = \frac{2\beta\omega}{(\omega_0^2 - \omega^2)}$

$\hookrightarrow$  resonance behavior

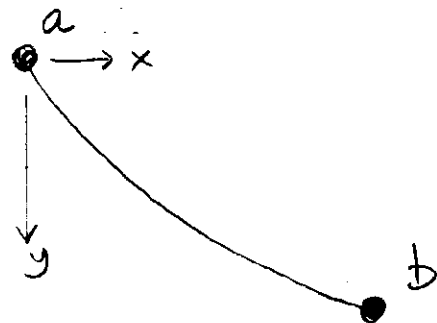


Arbitrary  $F(t)$

$x(t) = \left(\frac{1}{m\omega_1}\right) \int_{-\infty}^t F(t') \sin \omega_1(t-t') e^{-\beta(t-t')} dt'$

CHAPTER 6: Calculus of Variations

Typical kind of problem - Find the path (thru space) that minimizes time for rollercoaster to travel from a to b



$$dt = \frac{ds}{v}$$

$$\frac{1}{2}mv^2 = mgy \Rightarrow v = \sqrt{2gy}$$

$$t = \int \frac{1}{\sqrt{2gy}} [dx^2 + dy^2]^{\frac{1}{2}} = \text{minimum}$$

⇒ means nearby paths have nearly the same t

$$\delta \int \frac{1}{\sqrt{2gy}} [dx^2 + dy^2]^{\frac{1}{2}} = 0$$

e.g. use x as indep variable:  
y = y(x)

$$\delta \int_{x_1}^{x_2} \frac{[1 + (\frac{dy}{dx})^2]^{\frac{1}{2}}}{[2gy]^{\frac{1}{2}}} dx = 0$$

$$\delta \int_{x_1}^{x_2} F(y, y'; x) dx = 0$$

$$\frac{d}{dx} \frac{\partial F}{\partial y'} - \frac{\partial F}{\partial y} = 0$$

CHAP 7

HAMILTON'S PRINCIPLE:

$$\delta \int \mathcal{L} dt = \delta \int \mathcal{L}(q, \dot{q}, t) dt = 0$$

⇒

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{q}} - \frac{\partial \mathcal{L}}{\partial q} = 0$$

gives equations of motion  
Multiple coords ⇒ same eq for each.

$$\mathcal{L} = T - U$$

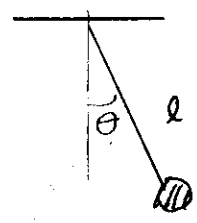
OK - good technique to find equations of motion in many complex systems

Definitions:  $p_k = \frac{\partial \mathcal{L}}{\partial \dot{q}_k}$  = generalized momentum

$Q_k = \frac{\partial \mathcal{L}}{\partial q_k}$  = generalized force  $\dot{p}_k = Q_k$

$$v = l\dot{\theta} \Rightarrow T = \frac{1}{2} m l^2 \dot{\theta}^2$$

$$U = mgl(1 - \cos\theta)$$



$$\mathcal{L} = \frac{1}{2} m l^2 \dot{\theta}^2 - mgl + mgl \cos\theta$$

$$P_{\theta} = m l^2 \dot{\theta} \quad Q_{\theta} = -mgl \sin\theta \quad m l^2 \ddot{\theta} = -mgl \sin\theta$$

angular mom.                      torque

CONSTANTS + CONSERVATION LAWS.

1) If  $\frac{\partial \mathcal{L}}{\partial q_k} = 0$  then  $\frac{d}{dt} P_k = 0 \Rightarrow P_k = \text{constant}$

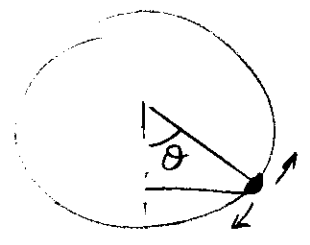
2) (a) Isolated system  $\Rightarrow$  all locations in space are equivalent  
 $\Rightarrow \delta \mathcal{L} = 0$  if  $\vec{r}_\alpha \rightarrow \vec{r}_\alpha + \vec{R}$  for all  $\alpha$   
 $\Rightarrow \vec{P}_{\text{TOT}} = \text{constant}$

(b) All directions in space are equivalent  $\vec{L}_{\text{TOT}} = \text{constant}$

3) If  $\frac{\partial \mathcal{L}}{\partial t} = 0$  then  $H \equiv \sum_k P_k \dot{q}_k - \mathcal{L} = \text{constant}$   
 $\hookrightarrow$  very often  $\hookrightarrow$  often T+U

• Good Example Problem:

Bead on a rotating hoop  $T = \frac{1}{2}$   
 $T = \frac{1}{2} m (R\dot{\theta})^2 + \frac{1}{2} m (R \sin\theta \omega)^2$



$$U = mgR(1 - \cos\theta)$$

$$\mathcal{L} = \frac{1}{2} m R^2 \dot{\theta}^2 + \frac{1}{2} m \omega^2 R^2 \sin^2\theta - mgR + mgR \cos\theta$$

- ① Eg. of mot. m
- ② H = constant.

$$H = \frac{1}{2} m R^2 \dot{\theta}^2 - \frac{1}{2} m \omega^2 R^2 \sin^2\theta + mgR - mgR \cos\theta = K$$

$\frac{1}{2} m \dot{x}^2 + U(x) = E$  - same form.

CHAP 8 TWO BODY CENTRAL FORCE PROBLEM.

$$\mathcal{L} = \frac{1}{2} M \dot{\vec{R}}^2 + \frac{1}{2} \mu \dot{\vec{r}}^2 - U(r)$$

↳ C.M. moves with const. vel.

$\vec{L}$  conserved  $\Rightarrow$  motion confined to a plane

$\Rightarrow$

$$\mathcal{L} = \frac{1}{2} \mu [\dot{r}^2 + r^2 \dot{\phi}^2] - U(r)$$

$\Rightarrow$

①  $\underline{l = \mu r^2 \dot{\phi} = \text{constant}}$

②  $\boxed{\mu \ddot{r} = \frac{l^2}{\mu r^3} - \frac{\partial U}{\partial r}}$

③  $\boxed{H = \frac{1}{2} \mu \dot{r}^2 + \frac{l^2}{2\mu r^2} + U(r) = E (\text{constant})}$

$$U_{\text{eff}} = U(r) + \frac{l^2}{2\mu r^2}$$

ORBIT EQUATION:  $u \equiv \frac{1}{r}$

$$\boxed{\frac{d^2 u}{d\phi^2} = -u - \left(\frac{\mu}{l^2}\right) \frac{1}{u^2} F\left(\frac{1}{u}\right)}$$

PLANETS:  $F = -\frac{K}{r^2} \Rightarrow \underline{\underline{r = \frac{c}{1 + \epsilon \cos \phi}}}$

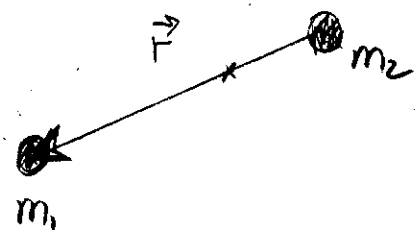
where

$$c = \frac{l^2}{K\mu}, \quad E = \frac{K^2 \mu}{2l^2} (\epsilon^2 - 1)$$

$\epsilon = 0 \Rightarrow$  circle ;  $0 < \epsilon < 1 \Rightarrow$  ellipse ;  $\epsilon = 1$  parabola

$\epsilon > 1 \Rightarrow$  hyperbola

Kepler's Laws follow from above.



## CHAPTER 9 - NON-INERTIAL FRAMES

### 1) Accelerating frames

$$m\ddot{\vec{r}} = \vec{F} - m\vec{A}$$

↳ "inertial force"

Constant acceleration  $\Rightarrow \vec{F}_{\text{inertial}}$  equivalent to uniform gravitational field

### • Tidal Forces

$$\vec{F}_{\text{tidal}} = GM_{\text{moon}} m \left[ \frac{\hat{d}_0}{d_0^2} - \frac{\hat{d}}{d^2} \right]$$

### 2) Rotating frames

$$\begin{aligned} m\vec{a} &= \vec{F} - 2m\vec{\Omega} \times \vec{v} - m\vec{\Omega} \times (\vec{\Omega} \times \vec{r}) \\ &= \vec{F} + \vec{F}_{\text{CORIOLIS}} + \vec{F}_{\text{CENTRIFUGAL}} \end{aligned}$$

$$\vec{F}_{\text{cent}} = m\Omega^2 \vec{\rho}$$

Near surface of earth  $\vec{g}_{\text{eff}} = \vec{g}_0 - \Omega^2 R_E \sin\theta \hat{\rho}$

$$\Rightarrow m\vec{a} = m\vec{g}_{\text{eff}} + \vec{F}' - 2m\vec{\Omega} \times \vec{v}$$

↳ other non-gravitational forces

### CORIOUS EFFECTS:

- air circulation around low pressure areas (hurricanes)
- projectile motion
- Foucault pendulum